

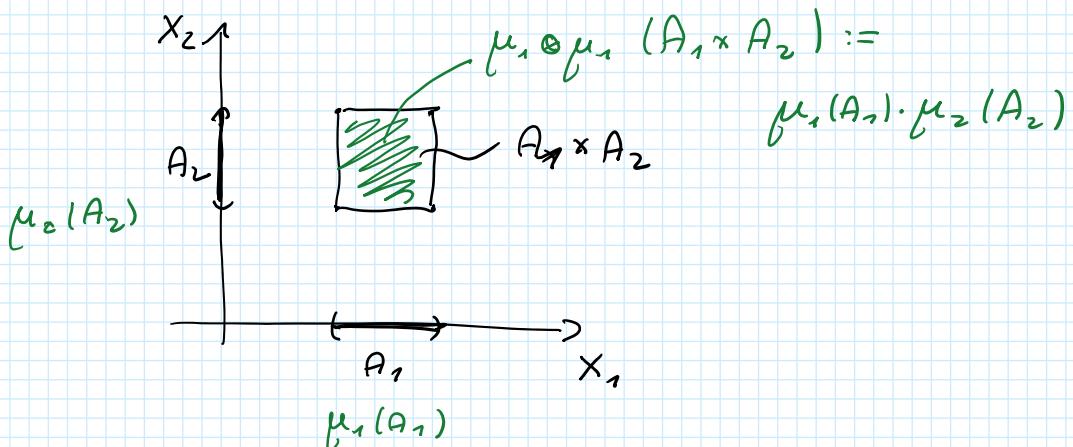
Addendum to integration theory: $\int \left(\int f(x, y) dx \right) dy = \int \left(\int f(x, y) dy \right) dx$

Product measures and Fubini's theorem

Def.: Let (X_1, \mathcal{A}_1) and (X_2, \mathcal{A}_2) be measurable spaces. Then $\mathcal{A}_1 \otimes \mathcal{A}_2$ denotes the σ -algebra on $X_1 \times X_2$ generated by sets of the form $A_1 \times A_2 \subset X_1 \times X_2$, with $A_1 \in \mathcal{A}_1$ and $A_2 \in \mathcal{A}_2$, the so called product σ -alg.

Example: Let $\mathcal{D}^n \subset \mathcal{P}(\mathbb{R}^n)$ be the Borel- σ -algebra.

Then $\mathcal{D}^n = \mathcal{D}^1 \otimes \dots \otimes \mathcal{D}^1$.



Thm: Let $(X_1, \mathcal{A}_1, \mu_1)$ and $(X_2, \mathcal{A}_2, \mu_2)$ be σ -finite measure spaces. There exists a unique measure μ on $\mathcal{A}_1 \otimes \mathcal{A}_2$ such that for all $A_1 \in \mathcal{A}_1, A_2 \in \mathcal{A}_2$

$$\mu(A_1 \times A_2) = \mu_1(A_1) \cdot \mu_2(A_2),$$

called the product measure and denoted by $\mu = \mu_1 \otimes \mu_2$.

Example: The Lebesgue-Borel measure

$$\lambda^n = \lambda^1 \otimes \cdots \otimes \lambda^1.$$

Theorem: Tonelli

Let $(X_1, \mathcal{A}_1, \mu_1)$ and $(X_2, \mathcal{A}_2, \mu_2)$ be σ -finite measure spaces. Let $f: X_1 \times X_2 \rightarrow [0, \infty]$ be $\mathcal{A}_1 \otimes \mathcal{A}_2$ -measurable. Then

$$\int_{X_1 \times X_2} f \, d(\mu_1 \otimes \mu_2) = \int_{X_1} \left(\int_{X_2} f_{x_1} \, d\mu_2 \right) \, d\mu_1,$$

$\underbrace{\quad}_{\text{Pct. of } x_1}$

$$= \int_{X_2} \left(\int_{X_1} f_{x_2} \, d\mu_1 \right) \, d\mu_2$$

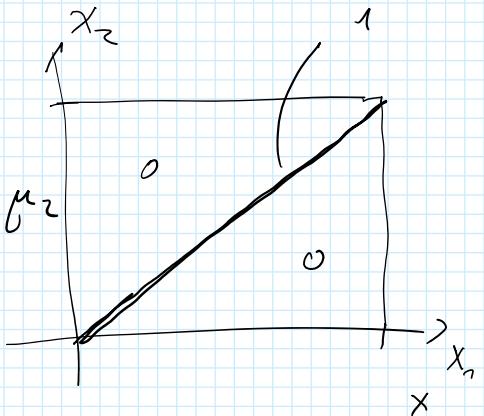
where $f_{x_1}: X_2 \rightarrow \overline{\mathbb{R}}$, $x_2 \mapsto f(x_1, x_2)$.

Ex.: $X_1 = X_2 = [0, 1]$, $\underline{\mu}_1 = \lambda^1$, $\mu_2 = \nu$ counting measure

$f: X_1 \times X_2 \rightarrow [0, \infty)$ (X_2, μ_2) is not σ -finite

$$(x_1, x_2) \mapsto f_{x_1, x_2} := \begin{cases} 1 & \text{if } x_1 = x_2 \\ 0 & \text{otherwise} \end{cases}$$

$$\int \underbrace{f(x_1, x_2)}_{f_{x_2}(x_1)} \, d\mu_1 = 0, \quad \int \left(\int f_{x_2} \, d\mu_1 \right) \, d\mu_2 = 0$$



$$\int f(x_1, x_2) d\mu_2 = 0 \cdot \underbrace{\mu_2(\chi_{\{f=0\}})}_{=0} + 1 \cdot \underbrace{\mu_2(\chi_{\{f=1\}})}_{2x_1, 5}$$

$\stackrel{!!}{f_{x_1}(x_2)}$

$$\begin{array}{c} 1 \\ \uparrow \\ \text{---} \\ x_1 \end{array} \quad = 0 + 1 = 1 \quad \int \left(\int f(x_1, x_2) d\mu_2 \right) d\mu_1 = 1$$

Thm.: Fubini

Let $(X_1, \mathcal{A}_1, \mu_1)$ and $(X_2, \mathcal{A}_2, \mu_2)$ be σ -fin. meas. spaces.

and $f: X_1 \times X_2 \rightarrow \mathbb{C}$ measurable.

(a) $\int_{X_1} \left(\int_{X_2} |f_{x_1}| d\mu_2 \right) d\mu_1 < \infty$ or if $\int_{X_2} \left(\int_{X_1} \dots d\mu_1 \right) d\mu_2 < \infty$

iff $f \in L^1(X_1 \times X_2, \mu_1 \otimes \mu_2)$.

(b) If $f \in L^1(X_1 \times X_2, \mu_1 \otimes \mu_2)$, then

$$\begin{aligned} \int_{X_1 \times X_2} f d(\mu_1 \otimes \mu_2) &= \int_{X_1} \left(\int_{X_2} f_{x_1} d\mu_2 \right) d\mu_1 \\ &= \int_{X_2} \left(\int_{X_1} f_{x_2} d\mu_1 \right) d\mu_2 \end{aligned}$$

Example: $X = \mathbb{N}$, $\mathcal{A} = \mathcal{P}(\mathbb{N})$, $\mu = v$ counting measure

$f: X \rightarrow \mathbb{R}$, $n \mapsto f_n$ (real sequence)

$$\int f d\mu = \sum_{n=1}^{\infty} f_n$$

$$\int_X f d\mu = \sum_{n=1}^{\infty} f_n$$

$g : X \times X \rightarrow \mathbb{R}$, $(n, m) \mapsto g_{n,m}$

$$\sum_n \sum_m g_{n,m} = ? \quad \text{if } \sum_m \sum_n |g_{n,m}| < \infty$$

$$\sum_m \sum_n |g_{n,m}| < \infty$$

$g_{n,m}$	$n=1$	2	3	4	5	
$m=1$	1	-1	0	0	0	.
2	0	1	-1	0	0	.
3	0	0	1	-1	0	.
4	0	0	0	1	-1	.
.

$$\sum_n g_{n,m} = 0 \quad \forall m$$

$$\sum_m \sum_n g_{n,m} = 0$$

$$\sum_m g_{n,m} = \begin{cases} 1 & \text{if } n=1 \\ 0 & \text{if } n>1 \end{cases}$$

$$\sum_n \sum_m g_{n,m} = 1$$

Preface on the two physics lectures

A physical theory is a (mathematical) model

for how (parts of) the physical world could work.

Physics is about

- inventing resp. discovering good „theories“
- collecting empirical data (experiments)
- comparing the empirical facts about our world with our theoretical worlds.

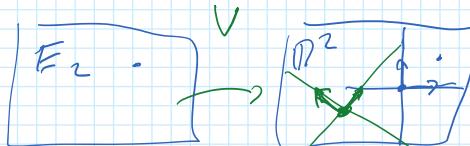
Quote from Democrit (400 BC):

"Apparently there is color, apparently sweetnes, apparently bitterness, while in reality there are only atoms and empty space."

Modern physics uses mathematical models because (at ~~that~~ least in good models) empirical predictions can be deduced by mathematical methods (computing, simulating, proving) in an unambiguous way.

Classical mechanics of point particles

Newtonian mechanics: Math. model for the motion of N particles (apples, planets, atoms, bullets, ...) in physical space and time.



Space: \mathbb{E}^3 three-dim. eucl. space,

\mathbb{R}^3 after choice of an origin and a ONB.

Time: \mathbb{E}^1 eucl. line, \mathbb{R} after choice of origin.

Configuration space of N point particles

$$q \in \mathbb{R}^{3N} = \underbrace{\mathbb{R}^3 \times \dots \times \mathbb{R}^3}_{N\text{-copies}} = \text{config. space}$$

$q = (q_1, q_2, \dots, q_n)$, $q_j \in \mathbb{R}^3$ position of j th particle.

In this model the "world" is completely specified by the positions of all particles at all times, i.e.

by a curve $\gamma: \mathbb{R} \rightarrow \mathbb{R}^{3N}$, $t \mapsto \gamma(t)$

in configuration space.

The physical "law" is a second order ODE for γ ,

Newton's law:

$$\ddot{\gamma} = M^{-1} \cdot F(t, \gamma, \dot{\gamma})$$

mass matrix

force field
 $\sim \dot{\gamma}(t)^2$

acceleration

Only the solutions to this ODE are possible worlds according to Newt. mech.

Assuming sufficient regul. of F , a unique solution is determined by specifying the positions $\gamma(t_0)$ and the velocities $\dot{\gamma}(t_0)$ at some time $t_0 \in \mathbb{R}$
 \Rightarrow predictions of the theory

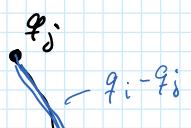
The explicit specification of M and F are also part of the law.

Example: Gravitating bodies: $M = \begin{pmatrix} m_1 & & 0 \\ & \ddots & \\ 0 & & m_N \end{pmatrix}$

m_i : mass of the i th body,

$$\|F_{j|i}\| \sim \frac{1}{\|q_i - q_j\|^2}$$

$$F_j(t, q, v) = \bar{F}_j(q) = G \sum_i \frac{m_i m_j (q_i - q_j)}{\|q_i - q_j\|^3}$$



$$\underline{\underline{F_j(t, q, v)}} = \underline{\underline{F_j(q)}} = G \sum_{i \neq j} \frac{m_i m_j (q_i - q_j)}{\|q_i - q_j\|^3}$$

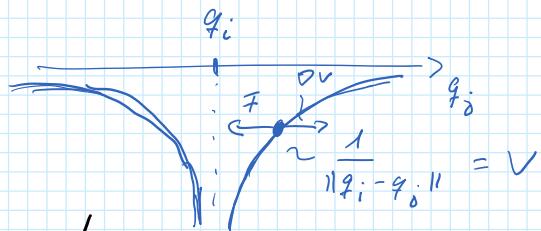
For $N=2$ one finds Kepler's ellipses as special solutions \Rightarrow Kepler's laws follow from Newtonian gravitation

The gravitational force is an example of a conservative force field, i.e. a force $\underline{\underline{F}}: \mathbb{R}^{3N} \rightarrow \mathbb{R}^{3N}$ that is the negative gradient of a scalar fct. $V: \mathbb{R}^{3N} \rightarrow \mathbb{R}$, the so called potential, $\underline{\underline{F}} = -\nabla V$.

Ex. 1: Check that the Newton pot. is

$$V(q) = -\frac{G}{2} \sum_{i \neq j} \frac{m_i m_j}{\|q_i - q_j\|}$$

\nwarrow double sum



For conservative forces, Newtonian mechanics displays „conservation of energy“: The fcts.

$$E: \mathbb{R}^{3N} \times \mathbb{R}^{3N} \rightarrow \mathbb{R}$$

↙ kinetic energy

$$E(q, \dot{q}) := \sum_{j=1}^N \frac{m_j}{2} \|\dot{q}_j\|^2 + \underline{\underline{V(q)}}$$

↙ pot. energy

is constant along solutions of $\ddot{q} = -m^{-1} \nabla V(q)$, i.e.

$$E(\gamma(t), \dot{\gamma}(t)) = E(\gamma(t_0), \dot{\gamma}(t_0)) \quad \forall t \in \mathbb{R}$$

Ex. 2: show this!

$$\begin{aligned}
\frac{d}{dt} E(\gamma(t), \dot{\gamma}(t)) &= \underbrace{\nabla_q E(\gamma(t), \dot{\gamma}(t))}_{n} \dot{\gamma}(t) + \underbrace{\nabla_v E(\gamma(t), \dot{\gamma}(t))}_{m_j} \dot{v}_j(t) \\
&= \langle \nabla V(\gamma(t)), \dot{\gamma}(t) \rangle_{\mathbb{R}^{3N}} + \sum_{j=1}^m \underbrace{\langle \nabla_{\dot{v}_j} V(\gamma(t)), \dot{v}_j(t) \rangle}_{\text{Lag}} \\
&\quad \nabla_{v_j} E(\gamma, \dot{\gamma}) \\
&= \langle \dot{\gamma}(t), \nabla V(\gamma(t)) \rangle_{\mathbb{R}^{3N}} + \langle \dot{\gamma}(t), \underbrace{M \dot{\gamma}(t)}_{\mathbb{R}^{3N}} \rangle_{\mathbb{R}^{3N}} \\
&= -\nabla V(\gamma(t)) \\
&= \text{Newton's eqn.}
\end{aligned}$$

Solutions stay on level sets of E !

If V is translation invariant, i.e.

$$V(q_1 + a, q_2 + a, \dots, q_N + a) = V(q_1, \dots, q_N) \quad \forall a \in \mathbb{R}^3$$

Then also total momentum

$$\begin{aligned}
\underline{P(q, v)} &= P(v) = \sum_{j=1}^N \underbrace{m_j v_j}_{=: p_j} \in \mathbb{R}^3 \\
&= \text{"momentum of part. } j \text{"}
\end{aligned}$$

is conserved.

If V is invariant under rotations of \mathbb{R}^3 ,

$$V(Rq_1, \dots, Rq_N) = V(q_1, \dots, q_N)$$

then angular momentum

$$L(q, v) := \sum_{j=1}^N m_j q_j \times v_j$$

is conserved. \Rightarrow symmetries lead to conservation laws.

As it is the case for any 2nd order ODE, one can write Newton's equation as a first order ODE on $\mathbb{R}^{6N} \Rightarrow$ Hamiltonian mechanics

Another very popular and useful formalism is the Lagrangian formulation of classical mechanics as a variational problem:

A Lagrange fct. is a fct

$$\mathcal{L}: \mathbb{R}^{3N} \times \mathbb{R}^{3N} \rightarrow \mathbb{R}, (\boldsymbol{q}, \boldsymbol{v}) \mapsto \underline{\mathcal{L}(\boldsymbol{q}, \boldsymbol{v})}$$

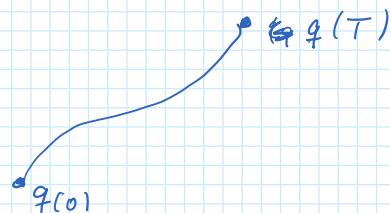
(e.g. $\underline{\mathcal{L}(\boldsymbol{q}, \boldsymbol{v})} = \sum_{j=1}^N \frac{m_j}{2} \|\dot{v}_j\|^2 - V(\boldsymbol{q})$)

Let $\Gamma := \{ \underline{\gamma} : C^2([t_0, T], \mathbb{R}^{3N}) \}$ the space of C^2 -paths in config.-space on time interval $[t_0, T]$.

The action of such a path is

$$S(\underline{\gamma}) := \int_0^T \mathcal{L}(\gamma(t), \dot{\gamma}(t)) dt$$

$$S: \Gamma \rightarrow \mathbb{R}$$



Then the principle of least action asserts that the physically possible paths are those for which S (when adding approp. const.) is critical, i.e.

Euler-Lagrange

$$\boxed{\mathcal{D}(S - \lambda \cdot H)|_x = 0}$$

(*)

Euler-Lagrange
equation

$$\text{As } D_S|_x \cdot h = D_{\dot{x}} \mathcal{L} \Big|_{(\dot{x}(T), \dot{\dot{x}}(T))} \cdot h(T) - D_{\dot{x}} \mathcal{L} \Big|_{(\dot{x}(0), \dot{\dot{x}}(0))} \cdot h(0)$$

$$+ \int_0^T \left\{ D_q \mathcal{L} \Big|_{(\dot{x}(t), \dot{\dot{x}}(t))} - \left(\frac{d}{dt} D_{\dot{x}} \mathcal{L} \Big|_{(\dot{x}(t), \dot{\dot{x}}(t))} \right) \right\} h(t) dt$$

a part of (*) is often (when h is only contrained at single points)

$$\boxed{D_q \mathcal{L} \Big|_{(\dot{x}(t), \dot{\dot{x}}(t))} - \frac{d}{dt} D_{\dot{x}} \mathcal{L} \Big|_{(\dot{x}(t), \dot{\dot{x}}(t))} = 0 \quad \forall t}$$

For $\mathcal{L} = \sum \frac{m_i}{2} \| \dot{x}_i \|^2 - V(x)$ these are exactly Newton's equations.

Hamiltonian mechanics

Phase space of N particles in \mathbb{R}^3 is

$P := \mathbb{R}^{6N} = \mathbb{R}^{3N} \times \mathbb{R}^{3N}$, $x \in P$ is of the form

$$x = (\underbrace{q_1, \dots, q_N}_{\text{positions}}, \underbrace{p_1, \dots, p_N}_{\text{momenta}})$$

(in general P is a symplectic space or manifold)

The canonical symplectic form on $P = \mathbb{R}^{6N}$ is

$$\omega : \mathbb{R}^{6N} \times \mathbb{R}^{6N} \rightarrow \mathbb{R} \quad (x, \dot{x}) \mapsto \dot{x} \cdot \nabla \mathcal{H}$$

$$f: \underline{\mathbb{R}^{6N}} \times \underline{\mathbb{R}^{6N}} \rightarrow \underline{\mathbb{R}}, (x_1, x_2) \mapsto \langle x_1, I x_2 \rangle_{\underline{\mathbb{R}^{6N}}}$$

with $I = \begin{pmatrix} 0 & \text{id}_{\mathbb{R}^{3N}} \\ -\text{id}_{\mathbb{R}^{3N}} & 0 \end{pmatrix}, I^T = -I$

The law of motion is a first order ODE on \underline{P}
where the vector field is the symplectic gradient
of a fct. $H: \underline{P} \rightarrow \underline{\mathbb{R}}$, the Hamiltonian:

$$\dot{\alpha} = I \nabla H(\alpha)$$

$\gamma(t)$
 t

now first order ODE
on phase space

$$\alpha: \underline{\mathbb{R}} \rightarrow P = \underline{\mathbb{R}^{6N}}$$

With $\alpha(t) := (\underline{Q}(t), \underline{P}(t))$ this reads

$$\begin{aligned} \begin{pmatrix} \dot{\underline{Q}}(t) \\ \dot{\underline{P}}(t) \end{pmatrix} &= \begin{pmatrix} 0 & \text{id} \\ -\text{id} & 0 \end{pmatrix} \begin{pmatrix} D_q H(\underline{Q}(t), \underline{P}(t)) \\ D_p H(\underline{Q}(t), \underline{P}(t)) \end{pmatrix} \\ &= \begin{pmatrix} D_p H(\underline{Q}(t), \underline{P}(t)) \\ -D_q H(\underline{Q}(t), \underline{P}(t)) \end{pmatrix} \end{aligned}$$

$$\text{For } \underline{H(q, p)} = \sum_{j=1}^N \frac{1}{2m_j} \|p_j\|^2 + V(q) = E(q, \underline{H^{-1}p})$$

one finds again Newton's eqnt.

$$\text{Let } \underline{\Phi}: \underline{P} \times \underline{P} \rightarrow \underline{P}, (t, x) \mapsto \alpha_x(t)$$

be the flow of a Hamiltonian system.

Then one has

"conservation of energy": $H \circ \underline{\Phi}_t = H \quad \forall t \in \mathbb{D}.$

"conservation of phase space volume":

$\underline{\Phi}_t^* \lambda = \lambda \quad \text{(i.e. } \lambda(\underline{\Phi}_t(A)) = \lambda(\underline{A}) \quad \forall \underline{A} \in \mathcal{D}(P) \text{)}$

Liouville's theorem (λ is Lebesgue-measure resp.
Liouville measure)

